# **RECOVERING SPACE CURVES BY MÖBIUS INVARIANTS\***

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**ABSTRACT:** One subgroup of the Möbius group in the Euclidean 3space  $\mathbb{R}^3$  is induced by the group of rotations on the unit 3-sphere in the Euclidean 4-space  $\mathbb{R}^4$  via a stereographic projection. We find two functions as Möbius invariants and they determine a Frenet space curve up to a transformation of considered group. We use these invariants to construct space curves and for that purpose we apply an algorithm in the computer system Mathematica.

**KEYWORDS:** Möbius group, Frenet space curves, Stereographic projection

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# 1 Introduction

The three dimensional unit sphere  $\mathbb{S}^3$ , centered at the origin O and embedded in the four-dimensional Euclidean space  $\mathbb{R}^4$ , is mapped via the stereographic projection  $\pi$  with a center at a point  $P \in \mathbb{S}^3$  onto an equatorial hyperplane  $\xi \subset \mathbb{R}^4$  through the origin O with a normal vector  $\overrightarrow{OP}$ . We consider the hyperplane  $\xi$  as a three-dimensional Euclidean space  $\mathbb{R}^3$ . The group of rigid motion on the sphere  $\mathbb{S}^3$  coincides with the group of rotations SO(4) in  $\mathbb{R}^4$  about the origin O. This group induces in  $\mathbb{R}^3$  via the stereographic projection  $\pi$  a subgroup  $M_0$  of the Möbius group Möb(3) in  $\mathbb{R}^3$ . For that group it is naturally to apply the algebra of quaternions  $\mathbb{H}$ , where  $\mathbb{H}_0 \subset \mathbb{H}$  is the set of pure quaternions (with zero real parts). If  $w \in \mathbb{H}$  is a quaternion then the norm of w is denoted by

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||w||, the quaternion  $\overline{w} \in \mathbb{H}$  is the conjugate quaternion of *w* and  $||w||^2 = w.\overline{w} = \overline{w}.w$ . The transformations from the group  $M_0$  can be represented by the quaternion formalism

$$F(q) = (A.q+B).(B.q+A)^{-1},$$

where  $A, B \in \mathbb{H}, q \in \mathbb{H}_0, ||A||^2 + ||B||^2 = 1$  and  $A.\overline{B} + B.\overline{A} = 0$ , identifying  $\mathbb{H}_0$  with  $\mathbb{R}^3$  (Theorem 3 in [2]).

It is well-known that the Euclidean curvature  $\varkappa$  and the Euclidean torsion  $\tau$  determine the space curves up to orientation preserving motions in  $\mathbb{R}^3$ . This fundamental theorem in the classical curve theory is extended to the orientation preserving similarity group in [1]. Recovering a space curve by a pair of real functions up to an orientation preserving motion or a similarity transformation is based on the mentioned theorems. The Euclidean geometry of curves on  $\mathbb{S}^3 \subset \mathbb{R}^4$  is carried naturally on  $\mathbb{R}^3$  by the stereographic projection  $\pi$  and it can be interpreted with respect to the subgroup  $M_0 \subset M\ddot{o}b(3)$  in  $\mathbb{R}^3$ . The differential geometry of curves with respect to the Möbius group is developed by Udo Hertrich-Jeromin in [7]. In the presented paper it is used a different approach. Relations between the Euclidean differential-geometric invariants of curves on the sphere  $\mathbb{S}^3$  and the corresponding curves in  $\mathbb{R}^3$ via the stereographic projection  $\pi$  are found. These relations determine invariants of curves with respect to the group  $M_0$ . They have an essential roll of recovering the curves in  $\mathbb{R}^3$  up to transformations from the considered group. The case of plane curves, recovering by one Möbuis invariant, is explored in [3] and [4].

#### 2 Preliminaries

Let  $K = O\vec{e}_1\vec{e}_2\vec{e}_3\vec{e}_4$  be a right-handed Cartesian coordinate system in  $\mathbb{R}^4$ . For any two vectors  $\boldsymbol{x} = (x^1, x^2, x^3, x^4)$  and  $\boldsymbol{y} = (y^1, y^2, y^3, y^4)$  in  $\mathbb{R}^4$  the canonical inner product  $\langle \boldsymbol{x}, \boldsymbol{y} \rangle$  is defined by  $\langle \boldsymbol{x}, \boldsymbol{y} \rangle = \sum_{i=1}^4 x^i y^i$ . We denote the norm of the vector  $\boldsymbol{x}$  by  $\|\boldsymbol{x}\| = \sqrt{\langle \boldsymbol{x}, \boldsymbol{x} \rangle}$ . The inner product and the norm of vectors in  $\mathbb{R}^3$  are determined in a

similar way.

Let  $\mathbb{S}^3 = \{(x^1, x^2, x^3, x^4) \in \mathbb{R}^4 | (x^1)^2 + (x^2)^2 + (x^3)^2 + (x^4)^2 = 1\}$ be the unit sphere in  $\mathbb{R}^4$ , centered at the origin *O*, and let  $l = (x^1, x^2, x^3, x^4), x^1 \neq 1$  be the position vector of an arbitrary point on  $\mathbb{S}^3$ , different from the pole P(1, 0, 0, 0). Denoting by  $(u^1, u^2, u^3)$  the coordinates of any position vector u of an arbitrary point in the equatorial hyperplane  $\xi \equiv \mathbb{R}^3$  with respect to the Cartesian coordinate system  $K' = O\vec{e}'_1\vec{e}'_2\vec{e}'_3$  in  $\mathbb{R}^3$ , where  $\vec{e}'_1 = \vec{e}_2$ ,  $\vec{e}'_2 = \vec{e}_3$ ,  $\vec{e}'_3 = \vec{e}_4$ , the stereographic map  $\pi : \mathbb{S}^3 \setminus \{P\} \to \mathbb{R}^3$  is given by the equalities  $u^i = \frac{x^{i+1}}{1-x^1}$ , i = 1, 2, 3. For the reverse map  $\pi^{-1}$  we have the equalities

$$x^{1} = \frac{(u^{1})^{2} + (u^{2})^{2} + (u^{3})^{2} - 1}{(u^{1})^{2} + (u^{2})^{2} + (u^{3})^{2} + 1},$$
  

$$x^{i} = \frac{2u^{i-1}}{(u^{1})^{2} + (u^{2})^{2} + (u^{3})^{2} + 1}, i = 2, 3, 4.$$

Making the infinity  $\infty$  correspond to the pole *P*, the stereographic projection  $\pi$  and the reverse map  $\pi^{-1}$  are bijective conformal maps. We consider the parametric equations of the sphere  $\mathbb{S}^3$  in the form

$$\boldsymbol{l}(u^{1}, u^{2}, u^{3}) = \left\{ \frac{\|\boldsymbol{u}\|^{2} - 1}{\|\boldsymbol{u}\|^{2} + 1}, \frac{2u^{1}}{\|\boldsymbol{u}\|^{2} + 1}, \frac{2u^{2}}{\|\boldsymbol{u}\|^{2} + 1}, \frac{2u^{3}}{\|\boldsymbol{u}\|^{2} + 1} \right\}.$$

The standard orthogonal tangential frame field  $\left\{\frac{\partial l}{\partial u^1}, \frac{\partial l}{\partial u^2}, \frac{\partial l}{\partial u^3}\right\}$  at any point of the sphere  $\mathbb{S}^3$ , except the point *P*, is given by

$$l_1 = \frac{\partial l}{\partial u^1} = \mu \cdot \left\{ u^1, \frac{1 - (u^1)^2 + (u^2)^2 + (u^3)^2}{2}, -u^1 u^2, -u^1 u^3 \right\}$$

$$l_{2} = \frac{\partial l}{\partial u^{2}} = \mu \cdot \left\{ u^{2}, -u^{1}u^{2}, \frac{1 + (u^{1})^{2} - (u^{2})^{2} + (u^{3})^{2}}{2}, -u^{2}u^{3} \right\}$$
  
$$l_{3} = \frac{\partial l}{\partial u^{3}} = \mu \cdot \left\{ u^{3}, -u^{1}u^{3}, -u^{2}u^{3}, \frac{1 + (u^{1})^{2} + (u^{2})^{2} - (u^{3})^{2}}{2} \right\},$$

where 
$$\mu = \frac{4}{(1+\|\boldsymbol{u}\|^2)^2}$$
 and  $l_1^2 = l_2^2 = l_3^2 = \mu$ .

A regular space curve in  $\mathbb{R}^3$  with a nowhere vanishing Euclidean curvature is called a Frenet space curve. For a given Frenet space curve  $c: u(s) = (u^1(s), u^2(s), u^3(s))$  in  $\mathbb{R}^3$ , parameterized by an arc-length parameter *s*, we denote by "/" the differentiation with respect to *s*. Applying the Frenet equations for the derivatives (see [5]), it is easy to prove the equalities

$$\langle \boldsymbol{u}', \boldsymbol{u}''' \rangle = - \langle \boldsymbol{u}'', \boldsymbol{u}'' \rangle = -\varkappa^2,$$

(1)

$$=\left(rac{oldsymbol{u}''^2}{2}
ight)'=arkalloldsymbol{arkall}',=-arkalloldsymbol{arkall}^2 au,$$

where  $\varkappa$  and  $\tau$  are the Euclidean curvature and the Euclidean torsion of *c*, respectively.

**Lemma 2.1.** Let v = v(s) be a vector function in  $\mathbb{R}^3$  with coordinate functions  $v^k(s) = \sum_{i,j=1}^3 \Gamma_{ij}^k u^{i'} u^{j'}$ , k = 1, 2, 3, where  $\Gamma_{ij}^k$  are the Christof-fel symbols for the sphere  $\mathbb{S}^3$  and  $u^i = u^i(s)$ , i = 1, 2, 3 are the coordinate functions of the unit speed curve c : u = u(s). Then

$$\boldsymbol{v} = \sqrt{\mu} \cdot \boldsymbol{u} - \mu \cdot \left(\frac{1}{\mu}\right)' \cdot \boldsymbol{u}'$$

and

$$< \boldsymbol{v}, \boldsymbol{u}^{\prime\prime\prime\prime}> = \varkappa^2 \mu \left(rac{1}{\mu}
ight)^\prime - \sqrt{\mu} \left(rac{\sqrt{\mu} + rac{\mu^2}{4} \left(rac{1}{\mu}
ight)^{\prime 2} - rac{\mu}{2} \left(rac{1}{\mu}
ight)^{\prime\prime}}{\sqrt{\mu}}
ight)^\prime,$$
  
 $\boldsymbol{v}^2 = \mu \boldsymbol{u}^2, < \boldsymbol{v}, \boldsymbol{u}^\prime> = -rac{\mu}{2} \left(rac{1}{\mu}
ight)^\prime, < \boldsymbol{u}^\prime imes \boldsymbol{u}^{\prime\prime\prime\prime}, \boldsymbol{v}> = \sqrt{\mu} < \boldsymbol{u} imes \boldsymbol{u}^\prime, \boldsymbol{u}^\prime\prime >^\prime,$ 

$$\langle \boldsymbol{v}, \boldsymbol{u}'' 
angle = -\sqrt{\mu} - rac{\mu^2}{4} \left(rac{1}{\mu}
ight)'^2 + rac{\mu}{2} \left(rac{1}{\mu}
ight)'',$$
  
 $\langle \boldsymbol{u}' imes \boldsymbol{u}'', \boldsymbol{v} 
angle = \mu \langle \boldsymbol{u} imes \boldsymbol{u}', \boldsymbol{u}'' 
angle.$ 

*Proof.* The Christoffel symbols  $\Gamma_{ij}^k$  for the sphere  $\mathbb{S}^3$  are given by

(2) 
$$\Gamma_{ij}^{k} = \Gamma_{ji}^{k} = \begin{cases} -\sqrt{\mu}.u^{j}, & k = i; \\ \sqrt{\mu}.u^{k}, & k \neq i = j; \\ 0, & k \neq i \neq j. \end{cases}$$

Hence,

$$\begin{split} \sum_{i,j=1}^{3} \Gamma_{ij}^{1} u^{i'} u^{j'} &= \Gamma_{11}^{1} (u^{1'})^{2} + 2\Gamma_{12}^{1} u^{1'} u^{2'} + 2\Gamma_{13}^{1} u^{1'} u^{3'} + \Gamma_{22}^{1} (u^{2'})^{2} + \\ &+ \Gamma_{33}^{1} (u^{3'})^{2} = -\sqrt{\mu} u^{1} (u^{1'})^{2} - 2\sqrt{\mu} u^{2} u^{1'} u^{2'} - 2\sqrt{\mu} u^{3} u^{1'} u^{3'} + \\ &+ \sqrt{\mu} u^{1} (u^{2'})^{2} + \sqrt{\mu} u^{1} (u^{3'})^{2} = -2\sqrt{\mu} u^{1'} (u^{1} u^{1'} + u^{2} u^{2'} + u^{3} u^{3'}) + \\ &+ \sqrt{\mu} u^{1} (\underbrace{(u^{1'})^{2} + (u^{2'})^{2} + (u^{3'})^{2}}_{=1}) = -\sqrt{\mu} (u^{2})' u^{1'} + \sqrt{\mu} u^{1}. \end{split}$$

Similarly, 
$$\sum_{i,j=1}^{3} \Gamma_{ij}^{2} u^{i'} u^{j'} = -\sqrt{\mu} (u^2)' u^{2'} + \sqrt{\mu} u^2$$
 and  $\sum_{i,j=1}^{3} \Gamma_{ij}^{3} u^{i'} u^{j'} = -\sqrt{\mu} (u^2)' u^{3'} + \sqrt{\mu} u^3$ . Since  $(u^2)' = ||u||^{2'} = \left(\frac{2}{\sqrt{\mu}} - 1\right)' = \sqrt{\mu} \left(\frac{1}{\mu}\right)'$ , we get the expression of the vector function  $v$ . The next scalar products

we get the expression of the vector function v. The next scalar products in the statement of the lemma are obtained by (1) and

$$< u, u' >= \frac{\sqrt{\mu}}{2} \left(\frac{1}{\mu}\right)'$$
 after the necessary differentiations.

The proofs of the next lemmas are routine, replacing the Christoffel symbols by (2).

**Lemma 2.2.** Let w = w(s) be a vector function in  $\mathbb{R}^3$  with coordinate functions  $w^k(s) = \sum_{i,j=1}^3 \Gamma_{ij}^k u^{i'} u^{j''}$ , k = 1, 2, 3, where  $\Gamma_{ij}^k$  are the Christof-fel symbols for the sphere  $\mathbb{S}^3$  and  $u^i = u^i(s)$ , i = 1, 2, 3 are the coordinate functions of the unit speed curve c : u = u(s). Then

$$\boldsymbol{w} = -\frac{\mu}{2} \left(\frac{1}{\mu}\right)' \cdot \boldsymbol{u}'' + \frac{1}{4} \left(4\sqrt{\mu} + \mu^2 \left(\frac{1}{\mu}\right)'^2 - 2\mu \left(\frac{1}{\mu}\right)''\right) \cdot \boldsymbol{u}'$$

and

$$< oldsymbol{w},oldsymbol{u}''> = -rac{\mu}{2}\left(rac{1}{\mu}
ight)'arkappa^2, < oldsymbol{w},oldsymbol{v}> = 0, \ < oldsymbol{w},oldsymbol{u}'> = - < oldsymbol{v},oldsymbol{u}''> = \sqrt{\mu} + rac{\mu^2}{4}\left(rac{1}{\mu}
ight)'^2 - rac{\mu}{2}\left(rac{1}{\mu}
ight)''.$$

**Lemma 2.3.** Let  $\eta = \eta(s)$  be a vector function in  $\mathbb{R}^3$  with coordinate functions  $\eta^k(s) = \sum_{i,j=1}^3 \Gamma_{ij}^k u^{i'} (u' \times u'')^j$ , k = 1, 2, 3, where  $\Gamma_{ij}^k$  are the Christoffel symbols for the sphere  $\mathbb{S}^3$ ,  $u^i = u^i(s)$ , i = 1, 2, 3 are the coordinate functions of the unit speed curve c : u = u(s) and  $(u' \times u'')^j$ , j = 1, 2, 3 are the coordinate functions of the vector product  $u' \times u''$  in  $\mathbb{R}^3$ . Then

$$\eta = -rac{\mu}{2}\left(rac{1}{\mu}
ight)'.(oldsymbol{u}' imes oldsymbol{u}'') - \sqrt{\mu} < oldsymbol{u} imes oldsymbol{u}', oldsymbol{u}'' > .oldsymbol{u}'.$$

and

$$< \eta, u'' > = < \eta, v > = 0, \ < \eta, u' > = - < u' imes u'', v > = -\sqrt{\mu} < u imes u', u'' > .$$

Let *c* be a Frenet space curve in  $\mathbb{R}^3$ . We denote by  $\gamma$  the stereographic pre-image of *c* on the sphere  $\mathbb{S}^3$ , i.e.  $\pi(\gamma) = c$ . Let  $l = l(\sigma)$ ,  $\sigma \in J$  be the vector parametric equation of  $\gamma$ , parameterized by an arc-length parameter  $\sigma$ . We have that  $l^2 = 1$ ,  $l'^2 = 1$  and  $\langle l, l' \rangle = 0$ ,  $\langle l, l'' \rangle = -1$ ,  $\langle l, l''' \rangle = 0$ . The vector field l + l'' is orthogonal to l and  $(l + l'')^2 = l''^2 - 1 \ge 0$ . A curve  $\gamma$  is said to be regular if  $(l + l'')^2 = l''^2 - 1 > 0$  for every  $\sigma \in J$ . A moving frame *ltnb* at any point of a regular curve  $\gamma$  on  $\mathbb{S}^3$  is introduced as follows: t = l',  $n = \frac{l + l''}{\|l + l''\|}$  and b is the unique vector, determined by the condition *ltnb* to be a right-handed orthonormal frame. The function  $\widetilde{\varkappa} = \|l + l''\| = \sqrt{l''^2 - 1} > 0$  is called a spherical curvature of  $\gamma$ . The vector field

$$n' + \widetilde{\varkappa}t = -rac{\widetilde{\varkappa}'}{\widetilde{\varkappa}^2}(l+l'') + rac{1+\widetilde{\varkappa}^2}{\widetilde{\varkappa}}l' + rac{1}{\widetilde{\varkappa}}l'''$$

has the direction of the vector  $\boldsymbol{b}$ . So, we can put  $\boldsymbol{n}' + \tilde{\varkappa} \boldsymbol{t} = \tilde{\tau} \boldsymbol{b}$ . The function  $\tilde{\tau}$  is called a spherical torsion of  $\gamma$ . For the moving frame ltnb we obtain the following Frenet type formulas

$$m{l}'=m{t}, \ m{t}'=-m{l}+\widetilde{arkappa}m{n}, \ m{n}'=-\widetilde{arkappa}m{t}+\widetilde{ au}m{b}, \ m{b}'=-\widetilde{ au}m{n}$$

If we denote by  $\nabla_t$  the covariant differentiation along the curve  $\gamma$  on  $\mathbb{S}^3$ then  $\nabla_t t = \tilde{\varkappa} n$ ,  $\nabla_t n = -\tilde{\varkappa} t + \tilde{\tau} b$  and  $\nabla_t b = -\tilde{\tau} n$ . Hence,  $\tilde{\varkappa} = \|\nabla_t t\|$ and  $\tilde{\tau} = -\langle \nabla_t b, n \rangle$  (see [6]).

#### 3 Main theorem

The fundamental theorem in the space curve theory states that the Euclidean curvature  $\varkappa$  and the Euclidean torsion  $\tau$  determine the space curve, parameterized by an arc-length parameter, up to a rigid motion in  $\mathbb{R}^3$ . Also, the spherical curvature  $\tilde{\varkappa}$  and the spherical torsion  $\tilde{\tau}$  define the curve, lying on the sphere  $\mathbb{S}^3 \subset \mathbb{R}^4$ , up to a rotation, preserving the sphere  $\mathbb{S}^3$ . Relations between the invariants  $\varkappa, \tau, \tilde{\varkappa}, \tilde{\tau}$  of the corresponding curves via the stereographic projection  $\pi$  will define any Frenet space curve up to transformation of the group  $M_0$ .

**Theorem 3.1.** Let c be a Frenet space curve in  $\mathbb{R}^3$ , parameterized by an arc-length parameter s, and let  $\gamma$  be its stereographic pre-image on

the sphere  $\mathbb{S}^3$ , i.e.  $\pi(\gamma) = c$ . If  $\varkappa = \varkappa(s)$  and  $\tau = \tau(s)$  are the Euclidean curvature and the Euclidean torsion of the curve c and  $\tilde{\varkappa} = \tilde{\varkappa}(s)$  and  $\tilde{\tau} = \tilde{\tau}(s)$  are the spherical curvature and the spherical torsion of the curve  $\gamma$  then

(3) 
$$\widetilde{\varkappa}^2 = \frac{1}{\mu}\varkappa^2 - 1 - \frac{3\mu}{4}\left(\frac{1}{\mu}\right)^2 + \left(\frac{1}{\mu}\right)^\prime$$

(4) 
$$\widetilde{\tau} = \frac{\cos\theta}{\mu\widetilde{\varkappa}}\varkappa\tau - \frac{\sin\theta}{\mu\widetilde{\varkappa}}\varkappa',$$

where

(5) 
$$\cos\theta = \frac{1}{\sqrt{\mu}\varkappa\widetilde{\varkappa}} \left(\varkappa^2 - \frac{1}{4}\mu^2 \left(\frac{1}{\mu}\right)'^2 + \frac{1}{2}\mu \left(\frac{1}{\mu}\right)'' - \sqrt{\mu}\right).$$

*Proof.* Let l = l(s) be a vector parametric equation of  $\gamma$  and  $\sigma = \sigma(s)$  be its arc-length function. If  $c : u(s) = (u^1(s), u^2(s), u^3(s))$  then  $\dot{l} = \frac{dl}{ds} = \sum_{i=1}^{3} (u^i)' l_i$ . The form  $d\sigma = \sqrt{\mu} ds$  is the linear element of the curve  $\gamma$  on the sphere  $\mathbb{S}^3$ . Thus, the unit tangent vector field is represented by  $t = \frac{1}{\sqrt{\mu}} \dot{l}$ . Applying the covariant differentiation  $\nabla_t$  we obtain that

(6) 
$$\nabla_{t} t = \sum_{k=1}^{3} \left( \frac{1}{\mu} (u^{k})^{\prime\prime} + \frac{1}{\mu} v^{k} + \frac{1}{2} \left( \frac{1}{\mu} \right)^{\prime} (u^{k})^{\prime} \right) l_{k},$$

where  $v^k = v^k(s)$ , k = 1, 2, 3 are the coordinate functions of the vector function v, defined in Lemma 2.1. Hence,  $\tilde{\varkappa}^2 =$ 

$$= \|\nabla_{t}t\|^{2} = \frac{1}{\mu}\varkappa^{2} + \frac{1}{\mu}v^{2} + \frac{\mu}{4}\left(\frac{1}{\mu}\right)^{\prime 2} + \frac{2}{\mu} < u^{\prime\prime}, v > + \left(\frac{1}{\mu}\right)^{\prime} < v, u^{\prime} >$$

and replacing with the equalities in Lemma 2.1 we get (3). The stereographic projection  $\pi$  and the reverse map  $\pi^{-1}$  are conformal maps and the differential  $\pi_*^{-1}$  of the reverse map  $\pi^{-1}$  is a similarity transformation between the corresponding tangent spaces with a coefficient  $\sqrt{\mu}$ .

If 
$$T = u', N = \frac{u''}{\varkappa}, B = T \times N = \frac{u' \times u''}{\varkappa}$$
 are the Frenet frame vec-

tor field along the curve c then  $\pi_*^{-1}(N) = \frac{1}{\varkappa} \sum_{i=1}^3 (u^i)'' l_i$  and  $\pi_*^{-1}(B) =$ 

 $\frac{1}{\varkappa}\sum_{i=1}^{3} (u' \times u'')^{i} \boldsymbol{l}_{i}. \text{ Vectors } \boldsymbol{\pi}_{*}^{-1}(\boldsymbol{N}) \text{ and } \boldsymbol{\pi}_{*}^{-1}(\boldsymbol{B}), \text{ lying in the plane, determined by vectors } \boldsymbol{n} \text{ and } \boldsymbol{b}, \text{ are orthogonal and let } \boldsymbol{\theta} = \angle(\boldsymbol{n}, \boldsymbol{\pi}_{*}^{-1}(\boldsymbol{N})).$ Then  $\boldsymbol{n} = \frac{\cos \boldsymbol{\theta}}{\sqrt{\mu}} \boldsymbol{\pi}_{*}^{-1}(\boldsymbol{N}) + \frac{\sin \boldsymbol{\theta}}{\sqrt{\mu}} \boldsymbol{\pi}_{*}^{-1}(\boldsymbol{B}).$  Hence,

$$\cos\theta = \frac{1}{\sqrt{\mu}} < \boldsymbol{n}, \pi_*^{-1}(\boldsymbol{N}) > = \frac{1}{\widetilde{\varkappa}\varkappa\sqrt{\mu}} < \nabla_{\boldsymbol{t}}\boldsymbol{t}, \sum_{i=1}^3 (u^i)''\boldsymbol{l}_i >$$

and

$$\sin \theta = \frac{1}{\sqrt{\mu}} < \boldsymbol{n}, \pi_*^{-1}(\boldsymbol{B}) > = \frac{1}{\widetilde{\varkappa} \varkappa \sqrt{\mu}} < \nabla_{\boldsymbol{t}} \boldsymbol{t}, \sum_{i=1}^3 (\boldsymbol{u}' \times \boldsymbol{u}'')^i \boldsymbol{l}_i > .$$

Using (6) we obtain that  $\cos \theta = \frac{1}{\widetilde{\varkappa} \varkappa \sqrt{\mu}} \left( u''^2 + \langle v, u'' \rangle \right) =$ 

 $=\frac{1}{\widetilde{\varkappa}\varkappa\sqrt{\mu}}\left(\varkappa^{2}+\langle \boldsymbol{v},\boldsymbol{u}''\rangle\right),$  whence applying Lemma 2.1 we get (5). Also,

(7) 
$$\sin \theta = \frac{1}{\widetilde{\varkappa} \varkappa \sqrt{\mu}} \langle v, u' \times u'' \rangle = \frac{1}{\widetilde{\varkappa} \varkappa} \langle u \times u', u'' \rangle.$$

Observing that  $\pi$  is an orientation preserving map in odd dimension, we have that  $\boldsymbol{b} = \frac{-\sin\theta}{\sqrt{\mu}} \pi_*^{-1}(\boldsymbol{N}) + \frac{\cos\theta}{\sqrt{\mu}} \pi_*^{-1}(\boldsymbol{B}) =$ 

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 $= -\frac{1}{\varkappa\sqrt{\mu}}\sum_{i=1}^{3} \left(\sin\theta(u^{i})'' - \cos\theta(u' \times u'')^{i}\right) \boldsymbol{l}_{i}.$  Therefore, using the covariant differentiation technique, we get

$$\nabla_{t}\boldsymbol{b} = \frac{1}{\sqrt{\mu}}\nabla_{l}\boldsymbol{b} = -\frac{1}{\sqrt{\mu}}\sum_{k=1}^{3} \left[\frac{1}{\varkappa\sqrt{\mu}}\left(\sin\theta(w^{k}+(u^{k})^{\prime\prime\prime})-\cos\theta(\eta^{k}+(u^{\prime}\times u^{\prime\prime\prime})^{k})\right) + \left(\frac{\sin\theta}{\varkappa\sqrt{\mu}}\right)^{\prime}(u^{k})^{\prime\prime} - \left(\frac{\cos\theta}{\varkappa\sqrt{\mu}}\right)^{\prime}(u^{\prime}\times u^{\prime\prime})^{k}\right]\boldsymbol{l}_{k}.$$

Thus,

$$\begin{split} \widetilde{\tau} &= - \langle \nabla_t b, n \rangle = -\frac{1}{\widetilde{\varkappa}} \langle \nabla_t b, \nabla_t t \rangle = \frac{\sqrt{\mu}}{\widetilde{\varkappa}} \left[ \frac{\sin \theta}{\varkappa \mu \sqrt{\mu}} \right] \\ & \left( \underbrace{\langle w, u'' \rangle + \langle u'', u'' \rangle + \langle w, v \rangle + \langle u''', v \rangle}_{\mathbf{A}} \right) + \frac{\sin \theta}{2\varkappa \sqrt{\mu}} \left( \frac{1}{\mu} \right)' \\ & \left( \underbrace{\langle w, u' \rangle + \langle u'', u'' \rangle}_{\mathbf{B}} \right) - \frac{\cos \theta}{\varkappa \mu \sqrt{\mu}} \right] \\ & \left( \underbrace{\langle \eta, u'' \rangle + \langle u' \times u''', u'' \rangle + \langle \eta, v \rangle + \langle u' \times u''', v \rangle}_{\mathbf{C}} \right) - \frac{\cos \theta}{\varkappa \sqrt{\mu}} \left( \frac{1}{\mu} \right)' \underbrace{\langle \eta, u' \rangle}_{\mathbf{D}} + \\ & + \frac{1}{\mu} \left( \frac{\sin \theta}{\varkappa \sqrt{\mu}} \right)' \left( \underbrace{\varkappa^2 + \langle u'', v \rangle}_{\mathbf{E}} \right) - \frac{1}{\mu} \left( \frac{\cos \theta}{\varkappa \sqrt{\mu}} \right)' \underbrace{\langle u' \times u'', v \rangle}_{\mathbf{F}} = \end{split}$$

$$= \frac{1}{\widetilde{\varkappa}} \left[ \frac{\sin \theta}{\varkappa \mu} . A + \frac{\sin \theta}{2\varkappa} \left( \frac{1}{\mu} \right)' . B - \frac{\cos \theta}{\varkappa \mu} . C - \frac{\cos \theta}{2\varkappa} \left( \frac{1}{\mu} \right)' . D + \frac{1}{\sqrt{\mu}} \left( \frac{\sin \theta}{\varkappa \sqrt{\mu}} \right)' . E - \frac{1}{\sqrt{\mu}} \left( \frac{\cos \theta}{\varkappa \sqrt{\mu}} \right)' . F. \right]$$

Applying the lemmas 2.1, 2.2, 2.3 and taking account of the equalities (1), (5), (7) we have that  $A = -\varkappa\varkappa' + \sqrt{\mu}(\varkappa\varkappa\cos\theta)'$ ,  $-B = E = \sqrt{\mu}\varkappa\varkappa\cos\theta$ ,  $C = -\varkappa^2\tau + \sqrt{\mu}(\varkappa\varkappa\sin\theta)'$ ,  $-D = F = \sqrt{\mu}\varkappa\varkappa\sin\theta$ . Replacing the found expressions in the equality above we obtain (4) after simple rearranges and simplifications.

### 4 Möbius invariants of space curves

It is obvious that the arc-length function  $\sigma = \sigma(s)$  is an invariant with respect to the group  $M_0$ . Let we denote by  $S_{\sigma}(s)$  the Schwarzian derivative of the function  $\sigma = \sigma(s)$ . Then we have that

$$S_{\sigma}(s) = \left(\frac{\sigma''}{\sigma'}\right)' - \frac{1}{2}\left(\frac{\sigma''}{\sigma'}\right)^2 = \frac{1}{2}\left[\frac{3}{4}\mu^2\left(\frac{1}{\mu}\right)'^2 - \mu\left(\frac{1}{\mu}\right)''\right].$$

If  $s = s(\sigma)$  is the reverse function of  $\sigma$  then we get  $0 = S_{\sigma \circ s}(\sigma) = S_{\sigma}(s) \left(\frac{ds}{d\sigma}\right)^2 + S_s(\sigma) \Rightarrow S_{\sigma}(s) = -\mu S_s(\sigma)$  applying the chain rule for the Schwarzian derivatives. The functions  $\tilde{\varkappa}$  and  $\tilde{\tau}$  from Theorem 3.1 can be expressed in terms of the arc-length parameter  $\sigma$ . We set  $\Re(\sigma) = \tilde{\varkappa}^2(\sigma)$ ,  $\mathfrak{T}(\sigma) = \tilde{\tau}(\sigma)$  and we obtain that

(8) 
$$\Re(\sigma) = \frac{1}{\mu}\varkappa^2 - 1 + 2S_s(\sigma)$$
$$\Im(\sigma) = \frac{1}{\mu\sqrt{\Re}} \left(\varkappa\tau\cos\theta - \sqrt{\mu}\frac{d\varkappa}{d\sigma}\sin\theta\right),$$
where  $\cos\theta = \frac{1}{\varkappa\sqrt{\mu\Re}} \left(\varkappa^2 + \mu S_s(\sigma) + \frac{1}{8\mu} \left(\frac{d\mu}{d\sigma}\right)^2 - \sqrt{\mu}\right).$ 

**Corollary 4.1.** The functions  $\mathfrak{K} = \mathfrak{K}(\sigma)$  and  $\mathfrak{T} = \mathfrak{T}(\sigma)$ , define by (8), are invariants under the group  $M_0$ .

Proof. The proof is omitted.

**Theorem 4.2** (Uniqueness theorem). Let  $I \subset \mathbb{R}$  be an open interval and let  $c_i : I \to \mathbb{R}^3$ , i = 1, 2 be two Frenet space curves, parameterized by the same arc-length parameter  $\sigma$  of their stereographic pre-images  $\gamma_i$ , i = 1, 2, respectively, on the sphere  $\mathbb{S}^3$  in  $\mathbb{R}^4$ . Assume that the curves  $c_1$  and  $c_2$  have the same invariants  $\Re_i = \Re_i(\sigma)$ ,  $\mathfrak{T}_i = \mathfrak{T}_i(\sigma)$ , i = 1, 2, defined by (8) for any  $\sigma \in I$ . Then there exists a transformation  $F \in M_0$  such that  $c_2 = F(c_1)$ .

*Proof.* Since  $\Re_1 = \Re_2$  and  $\mathfrak{T}_1 = \mathfrak{T}_2$  then the spherical stereographic preimages  $\gamma_1$  and  $\gamma_2$  of the curves  $c_1$  and  $c_2$ , respectively, via the stereographic projection  $\pi$ , have the same spherical curvatures and torsions. Therefore, there exists a rotation  $\rho \in SO(4)$  such that  $\gamma_2 = \rho(\gamma_1)$ . Hence,  $c_2 = \pi(c_1) = \pi \circ \rho(\gamma_1) = \pi \circ \rho \circ \pi^{-1}(c_1)$ , where  $F = \pi \circ \rho \circ \pi^{-1} \in M_0$ and the proof is completed.

**Theorem 4.3** (Existence theorem). Let  $f: I \to \mathbb{R}$ , f > 0 and  $g: I \to \mathbb{R}$  be given  $C^{\infty}$ -functions, defined on the same interval  $I \subset \mathbb{R}$ . Let  $c_0 \in \mathbb{R}^3$  and  $e_1^0$ ,  $e_2^0$ ,  $e_3^0$  be a right-handed orthonormal frame at  $c_0$  in the Euclidean space  $\mathbb{R}^3$ . There exists a unique Frenet space curve  $c: I \to \mathbb{R}^3$  which satisfies the conditions:

(a) there exists  $\sigma_0 \in I$ , such that  $c(\sigma_0) = c_0$ , and the Frenet frame of c at  $c_0$  is  $e_1^0, e_2^0, e_3^0$ ;

(b) for any  $\sigma \in I \ \Re(\sigma) = f^2(\sigma)$  and  $\mathfrak{T}(\sigma) = g(\sigma)$ .

*Proof.* Let  $l_0 = \pi^{-1}(c_0)$  and  $t_0 = \pi^{-1}_*(e_1^0)$ ,  $n_0 = \pi^{-1}_*(e_2^0)$ ,  $b_0 = \pi^{-1}_*(e_3^0)$ . Let us consider a matrix-valued function

 $\mathscr{E}(\sigma) = (\boldsymbol{l}(\sigma), \boldsymbol{t}(\sigma), \boldsymbol{n}(\sigma), \boldsymbol{b}(\sigma))^T$ . Solving the system of first order linear differential equations

$$\frac{d}{d\sigma}\mathcal{E} = \mathscr{A}(\sigma)\mathcal{E}, \text{ with a given matrix} \begin{pmatrix} 0 & 1 & 0 & 0\\ -1 & 0 & f & 0\\ 0 & -f & 0 & g\\ 0 & 0 & -g & 0 \end{pmatrix} \text{ and ini-}$$

tial conditions  $l_0$ ,  $t_0$ ,  $n_0$ ,  $b_0$  we obtain a unique solution  $\mathscr{E} = \mathscr{E}(\sigma)$ , determined for all  $\sigma \in I$  and  $\mathscr{E}(\sigma_0) = (l_0, t_0, n_0, b_0)^T$  for some  $\sigma_0 \in I$ . It is routine to prove that the matrix  $\mathscr{E}$  is orthogonal. This means that the vectors  $l(\sigma)$ ,  $t(\sigma)$ ,  $n(\sigma)$ ,  $b(\sigma)$  form an orthogonal frame in  $\mathbb{R}^4$  for any  $\sigma \in I$ . Let  $\gamma$  be a spherical curve, defined by the vector function  $l = l(\sigma)$  and let  $c = \pi(\gamma)$ . It is clear that the conditions (*a*) and (*b*) in the statement of the theorem are fulfilled for the space curve *c*.

The proof of the last theorem give us an algorithm of recovering space curves up to a transformation from the group  $M_0$ .

**Algorithm.** Recovering space curves by two functions  $f = f(\sigma) > 0$ and  $g = g(\sigma)$  for any  $\sigma \in I$ .

- 1. Choose initial conditions:  $c_0$ ,  $e_1^0$ ,  $e_2^0$ ,  $e_3^0$ ;
- 2. Find the vectors  $\boldsymbol{l}_0 = \pi^{-1}(\boldsymbol{c}_0), \, \boldsymbol{t}_0 = \pi^{-1}_*(\boldsymbol{e}_1^0), \, \boldsymbol{n}_0 = \pi^{-1}_*(\boldsymbol{e}_2^0), \, \boldsymbol{b}_0 = \pi^{-1}_*(\boldsymbol{e}_3^0);$

3. Solve the differential equation  $\frac{d}{d\sigma} \mathcal{E} = \mathscr{A}(\sigma)\mathcal{E}$ , where  $\begin{pmatrix} 0 & 1 & 0 & 0 \\ -1 & 0 & f & 0 \\ 0 & -f & 0 & g \\ 0 & 0 & -g & 0 \end{pmatrix}$ , and initial conditions, determined in Step 1., for  $\mathcal{E}(\sigma_0) = (\mathbf{l}_0, \mathbf{t}_0, \mathbf{n}_0, \mathbf{b}_0)^T$ ;

4. The curves  $\gamma$ , with vector function  $\mathbf{l} = \mathbf{l}(\sigma)$ , and  $c = \pi(\gamma)$  are found.

The next two examples illustrate the considered conformal invariants of a space curve in  $\mathbb{R}^3$ . For the calculations and visualizations we use the computer system Mathematica.

Example 4.4. Let

$$c: \boldsymbol{u}(\boldsymbol{\sigma}) = \left\{ \cos\left(\sqrt{2}\tan\left(\frac{\boldsymbol{\sigma}}{2}\right)\right), \sin\left(\sqrt{2}\tan\left(\frac{\boldsymbol{\sigma}}{2}\right)\right), \sqrt{2}\tan\left(\frac{\boldsymbol{\sigma}}{2}\right) \right\}$$

be a helix in  $\mathbb{R}^3$ , parameterized by an arc-length parameter  $\sigma$  of its spherical stereographic pre-image. Then  $\Re(\sigma) = \frac{1}{4\cos^4(\frac{\sigma}{2})}$  and  $\Im(\sigma) =$ 

 $\cos \sigma$ 

 $1 + \cos \sigma$ 

**Example 4.5.** Let  $\Re(\sigma) = \sigma^2$ ,  $\mathfrak{T}(\sigma) = 0.6\sigma$ ,  $c_0 = (0,0,0)$ ,  $e_1^0 = (0,0,1)$ ,  $e_2^0 = (0,1,0)$ . Then, applying the algorithm above, where  $f(\sigma) = \sigma$ , and  $g(\sigma) = 0.6\sigma$ , we obtain the Frenet space curve, depicted in Fig.1.



Fig. 1: 
$$\Re(\sigma) = \sigma^2$$
,  $\Im(\sigma) = 0.6\sigma$ 

#### **REFERENCES:**

- Encheva R. and Georgiev G., Similar Frenet Curves, *Result. Math.*, 55 no. 3-4 (2009) 359-372.
- [2] Encheva, R.P., Möbius transformations induced by rotations on the threesphere, MATTEX 2016, *Conference proceedings*, vol. **1** (2016) 43-50.
- [3] Encheva R., Recovering Plane Curves by One of Their Conformal Invariants, *Proceedings* vol. 53, book 6.1, Mathematics, Informatics and Phisics, Ruse, (2014) 22-27.
- [4] Encheva R., Family of Plane Curves in the Extended Gauss Plane Generated by One Function, Wolfram Demonstrations Project, Published: July 8, 2013 http://demonstrations.wolfram.com/FamilyOfPlaneCurvesInTheExtendedGaussPlaneGeneratedByOneFunc/
- [5] Künel W., Differential geometry: Curves-Surfaces-Manifolds, Second Edition, ISBN: 0-8218-3988-8, AMS, 2006.
- [6] Tazawa Y., Curves and surfaces in the three dimensional sphere placed in the space of quaternions, *Innovation in mathematics, Transactions on Engineering Sciences*, vol. 15 (1997) 459-466.
- [7] Udo Hertrich-Jeromin, Introduction to Möbius Differential Geometry, London Math. Soc. L. N. Series 300, Cambridge University Press, Cambridge, 2003.

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